

The Millennium-II Simulation:

Overview and database-enabled science

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with thanks to: Gerard Lemson, Volker Springel, Simon White, Adrian Jenkins



Millennium Workshop
Garching
17 December 2012

Extending the Millennium Simulation's reach to lower masses
(see Raul's talk for extending it to higher masses / rarer objects)

Simulation	Millennium Springel et al. 2005	Millennium-II MBK et al. 2009
N_p	2160^3	2160^3
Box Size	685 Mpc	137 Mpc
Mass Resolution	$1.2 \times 10^9 M_{\text{sun}}$	$9.4 \times 10^6 M_{\text{sun}}$
Force Resolution	6.85 kpc	1.37 kpc

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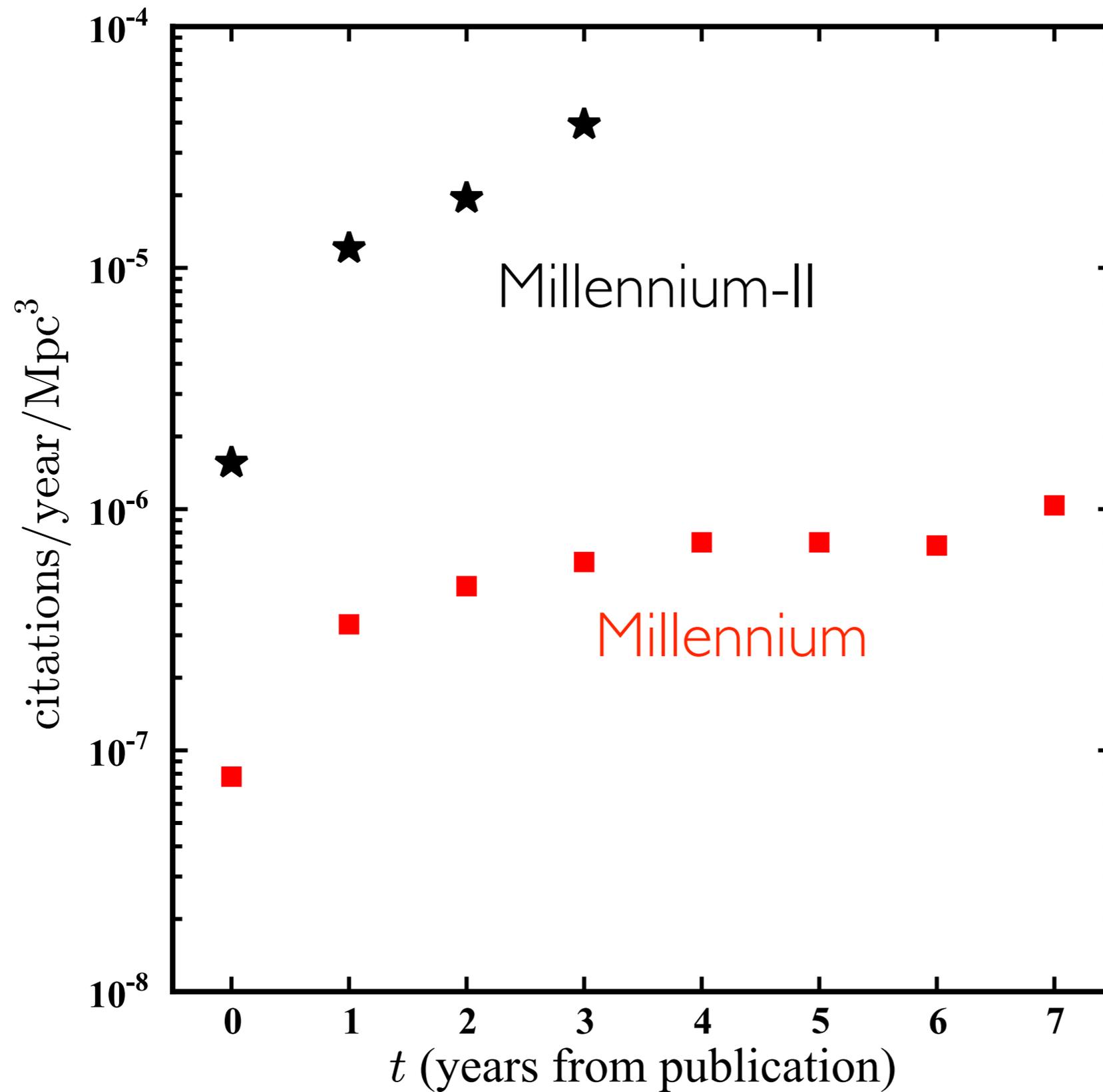
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# of citations	~1300 ~1500	~125 ~190
# of citations/Gpc³	~4050 ~4700	~48000 ~74000

(as of August 2011, **today**)

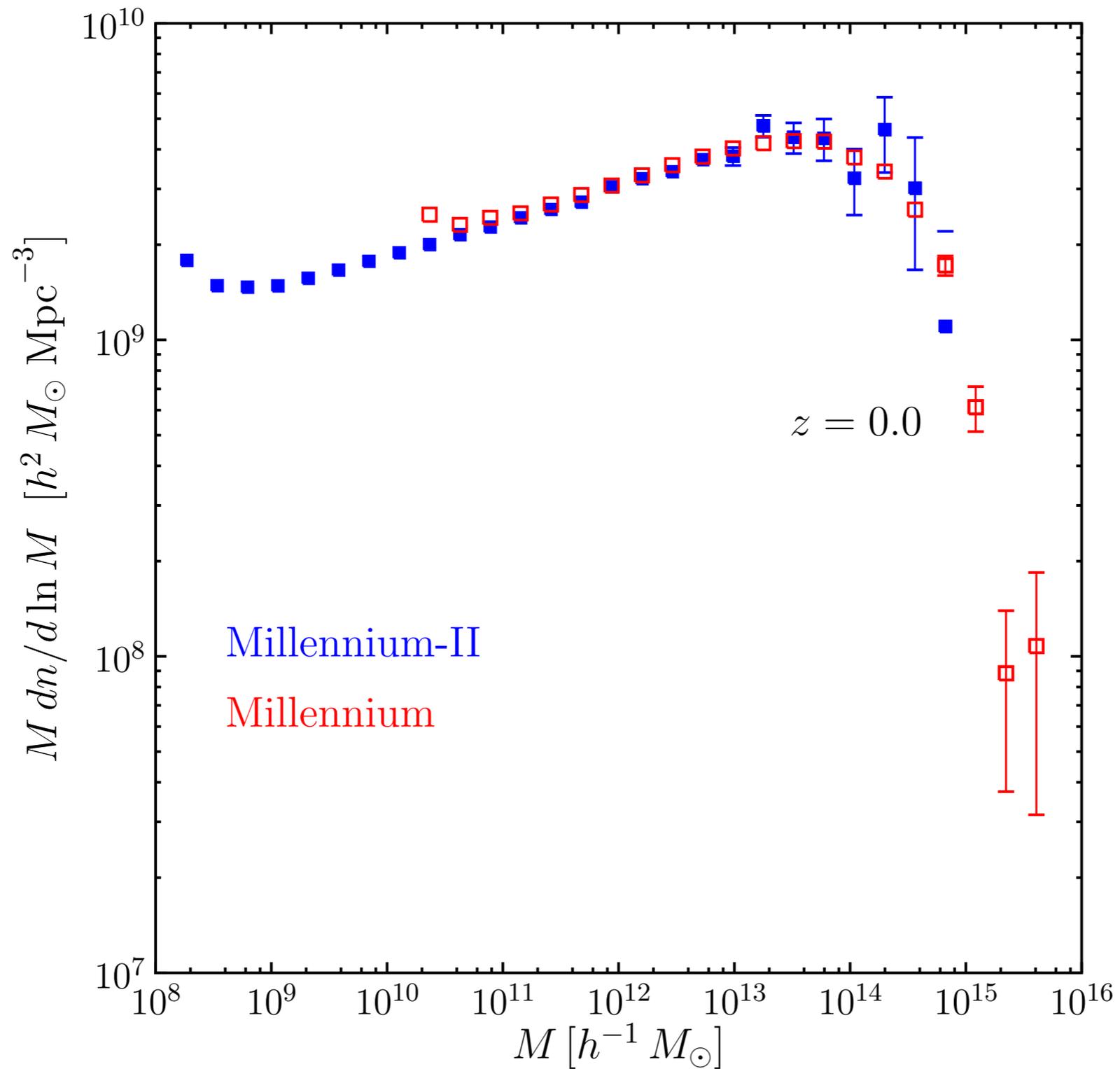
Cosmic citation rate density



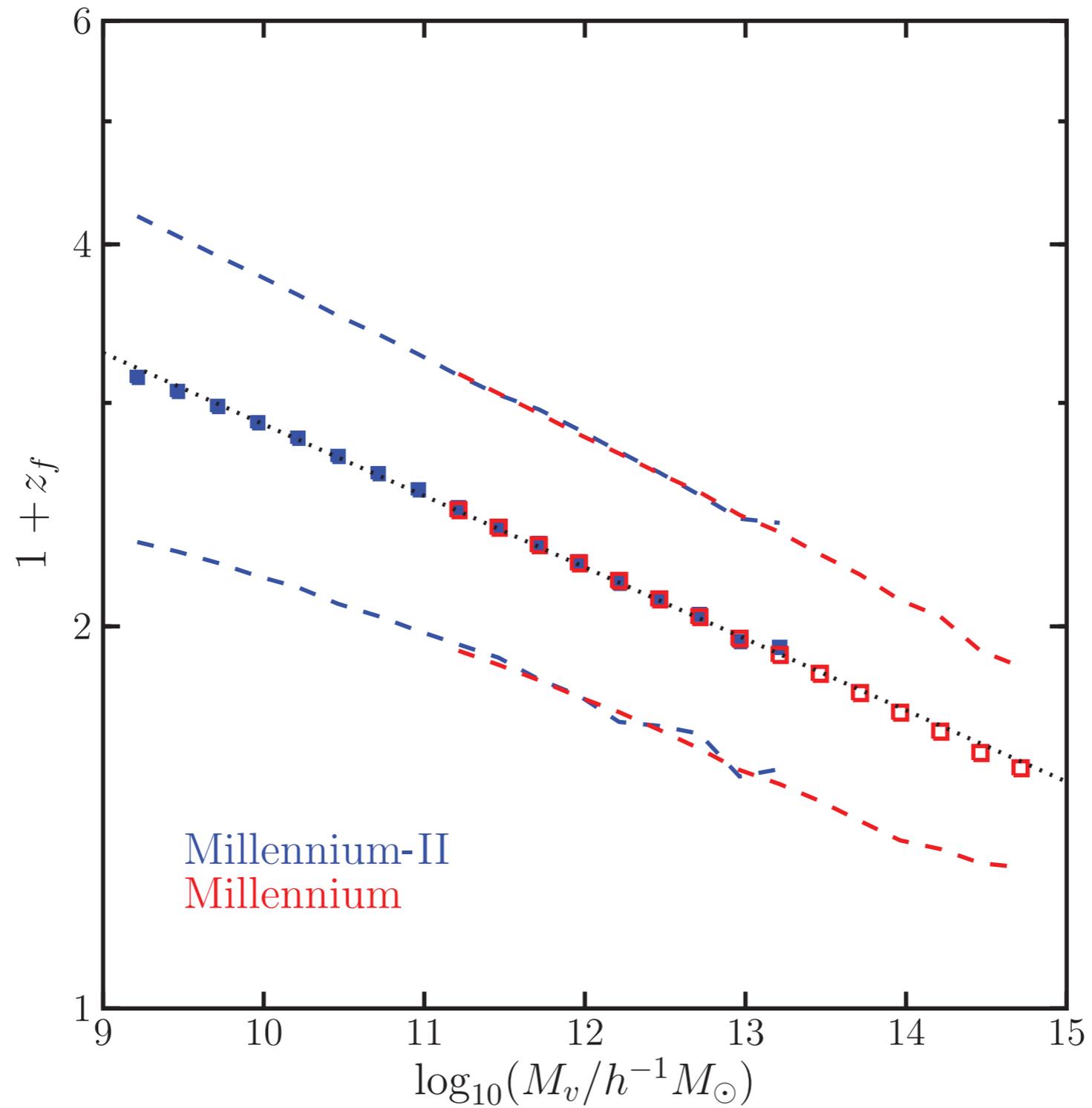
Primary Science Goals for Millennium-II

- Statistical understanding of the assembly, properties, and substructure of galaxy-mass dark matter halos
 - ▶ Needed for resolving the Milky Way and its bright satellites directly: $N_p = 10^5$
 - ▶ Good statistics: need > 1000 halos
 - ▶ Together, these set minimum requirement for MS-II (> 100 Mpc, $M_p < 10^7 M_{\text{sun}}$)
 - ▶ *Bonus*: use same initial conditions as run from which Aquarius halos were selected \Rightarrow Aquarius halos are present in MS-II as well
- Complement and extend Millennium Simulation in this regime
 - ▶ where does numerical resolution become an issue?
 - ▶ what changes are needed in galaxy formation models (c.f. Qi, Bruno)?
 - ▶ make predictions about satellite / galaxy pairs down to SMC scale

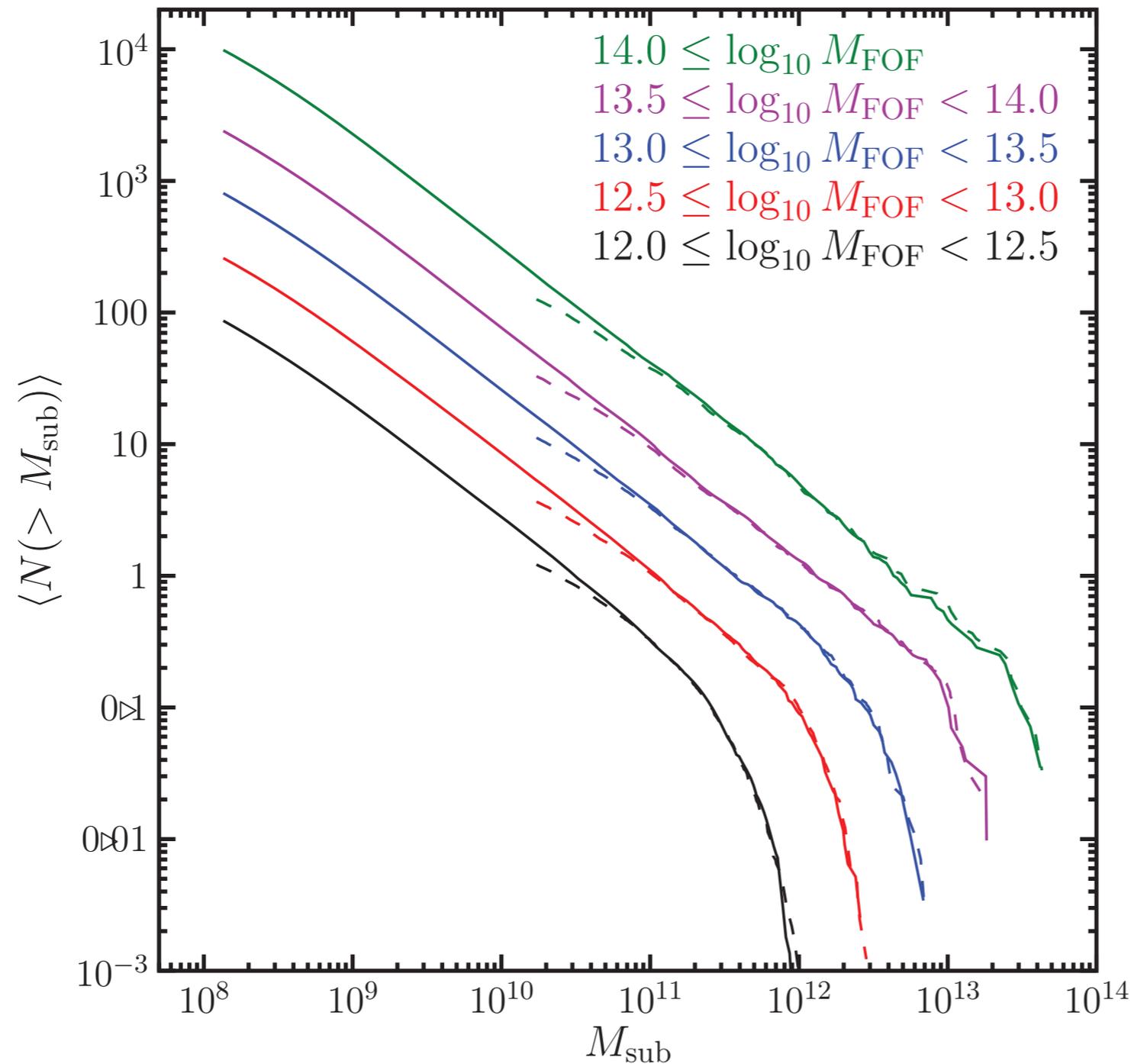
$z=0$ FOF mass function



Half-mass formation times

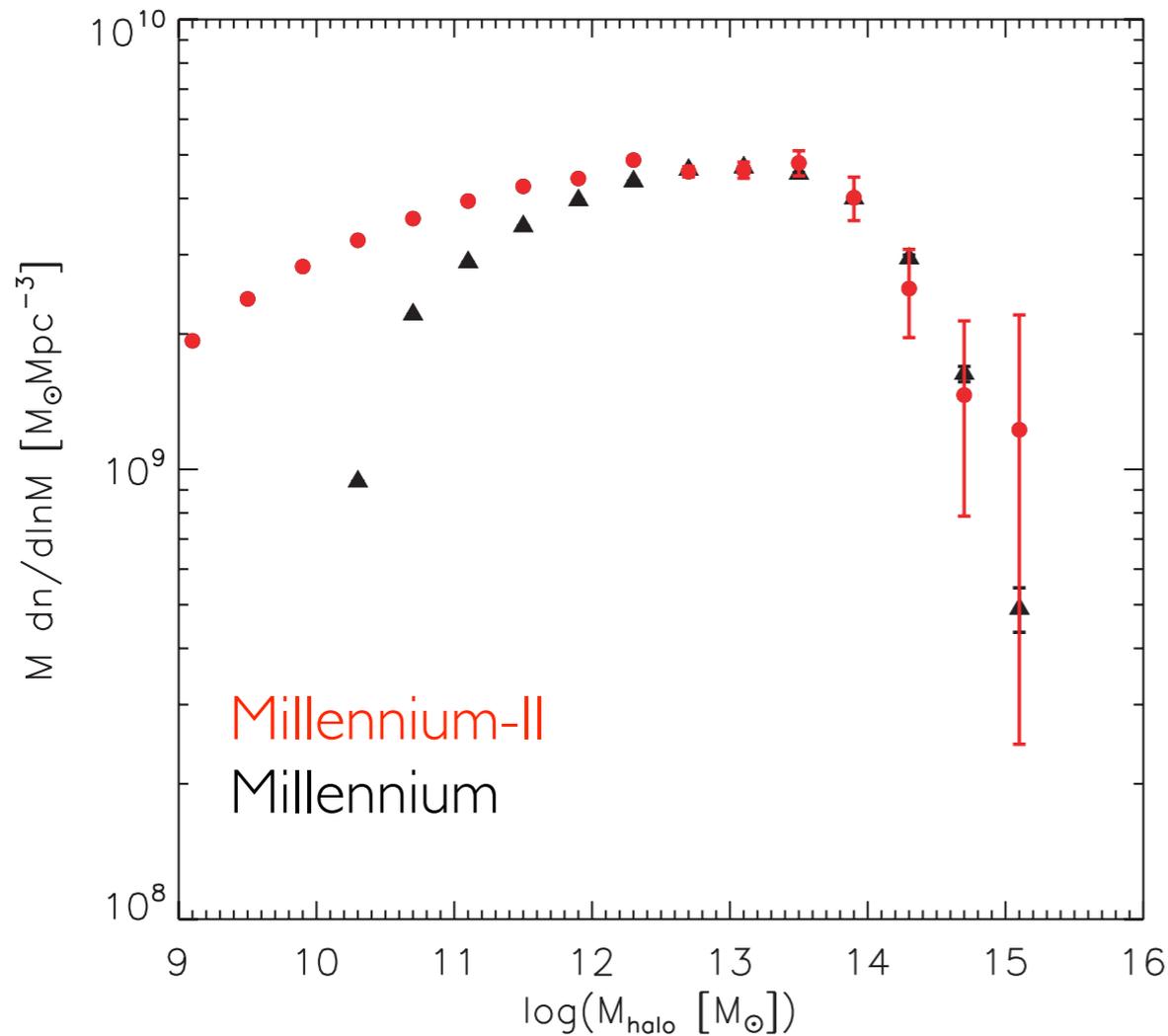


Subhalo abundance



Solid: MS-II
Dashed: mini-MS-II
(resolution of Millennium)

Resolution and semi-analytics

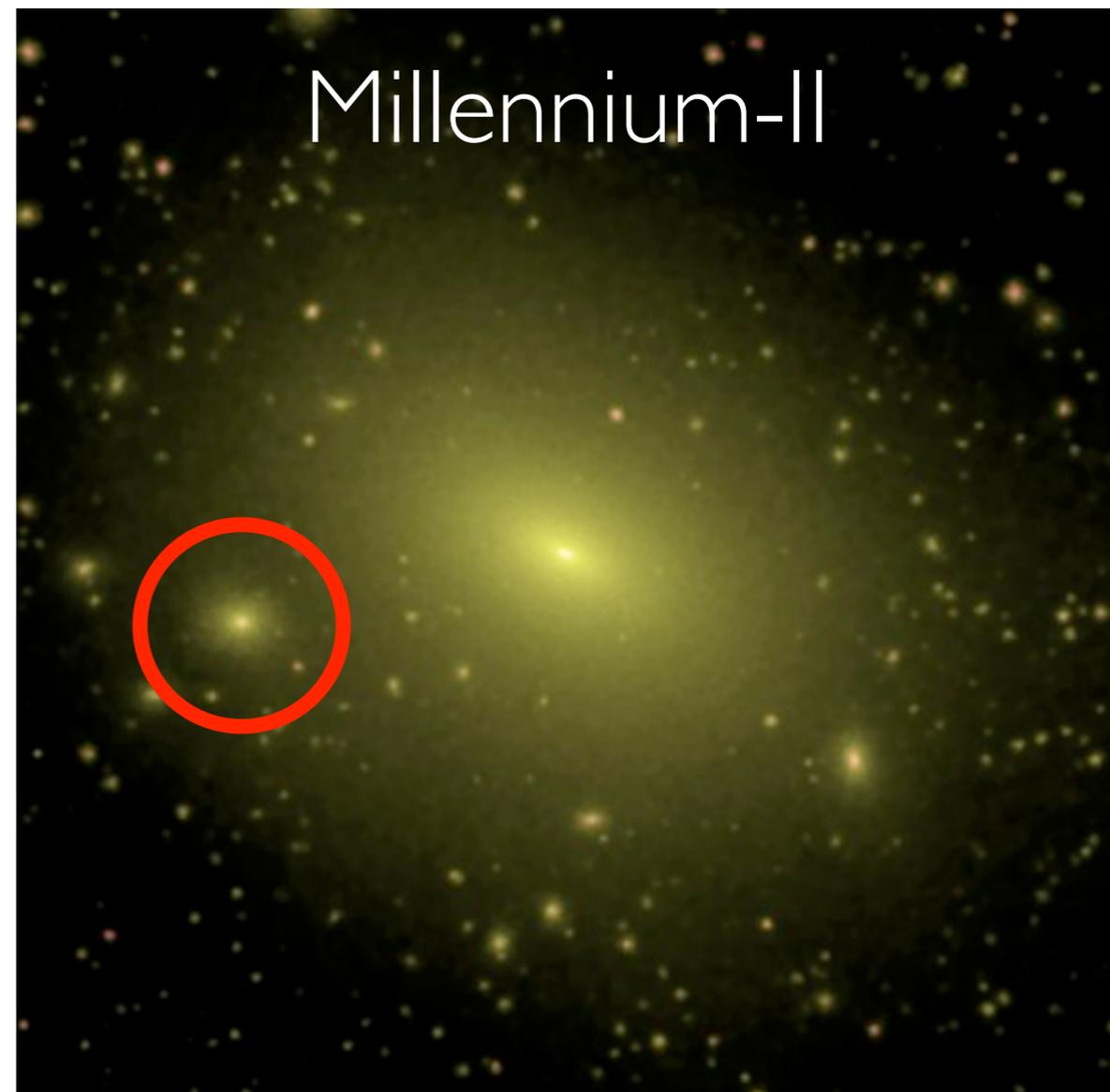
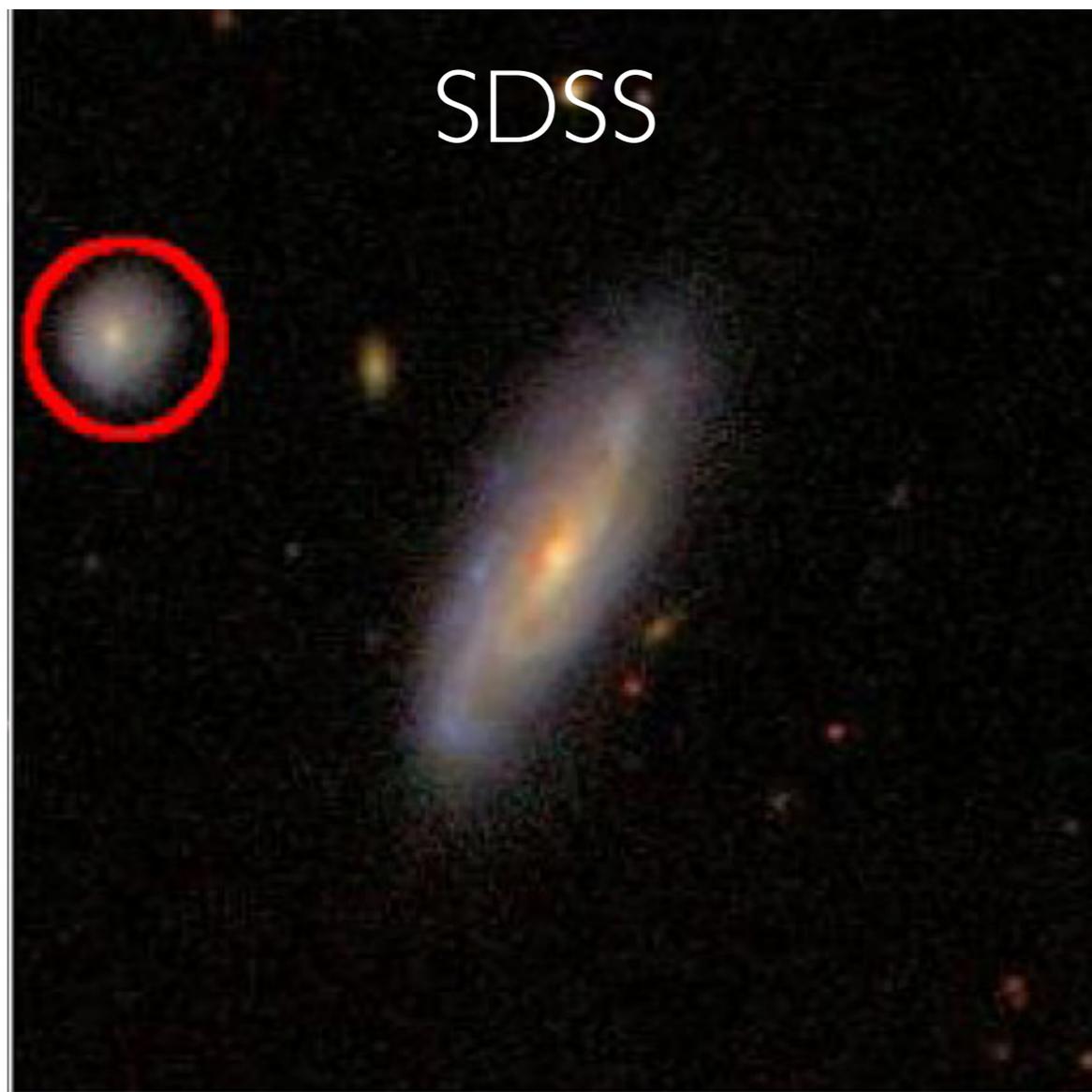


Guo, White, Li, MBK 2010

- Mass function based on infall mass indicates mass functions are complete only for **~2000** particles and greater (!!!!!)
- Numerical suppression of halo mass function leads to better, but artificial, agreement between galaxy luminosity function and halo mass function at faint end
- With Millennium + Millennium-II: can resolve all halos having $\log_{10}(M/M_{\text{sun}}) > 10$, i.e., full galaxy population observable outside of the Local Group

Satellite kinematics in the nearby Universe

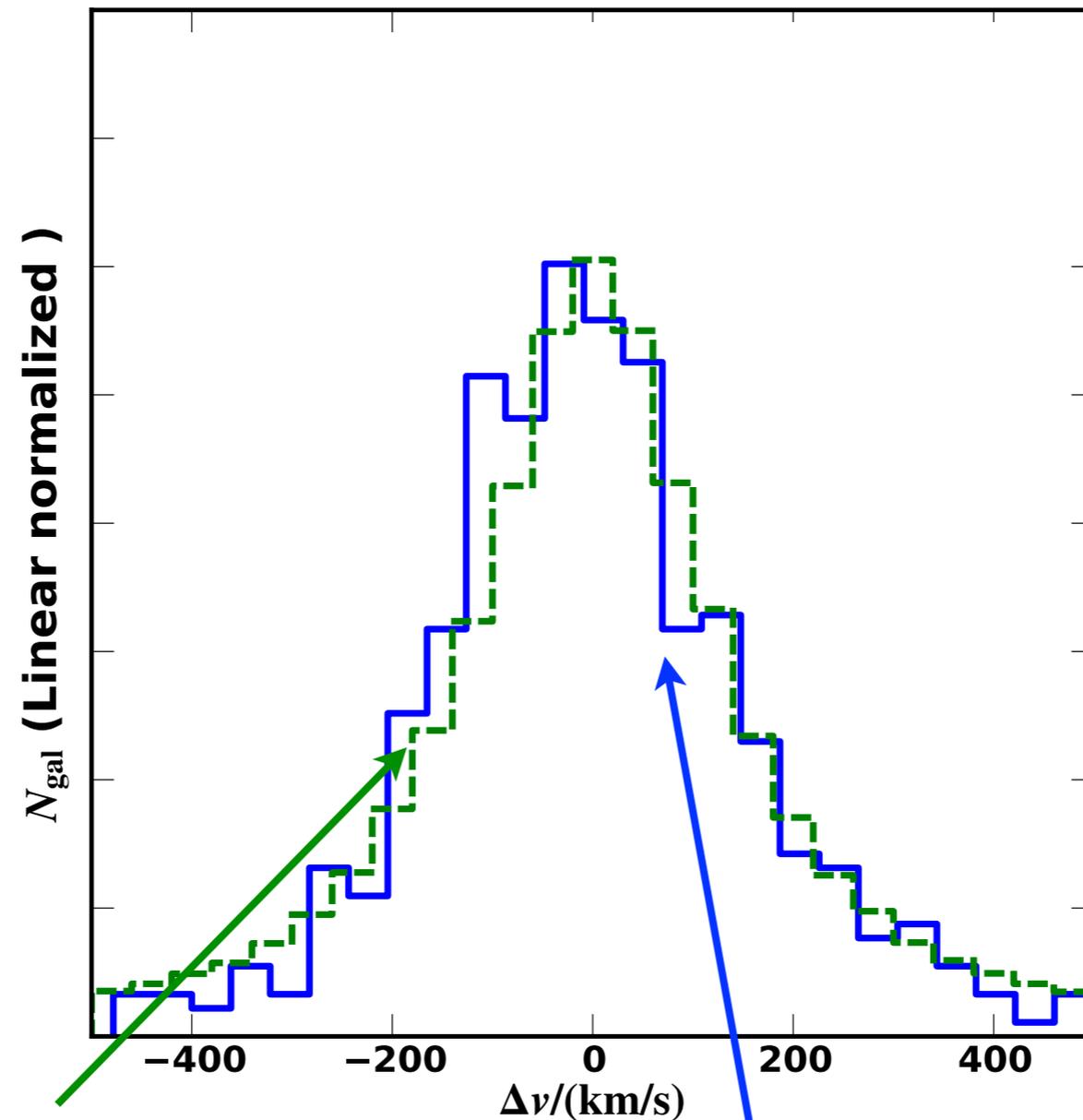
SDSS: spectroscopically complete to $M_r = -18.3$ (\sim LMC) within $z=0.034$; this volume matches Millennium-II very well.



Use MS-II to maximize purity / completeness of satellite-host pairs

Satellite kinematics in the nearby Universe

Pair-wise velocity dispersion and radial distribution of satellites match MS-II + abundance matching

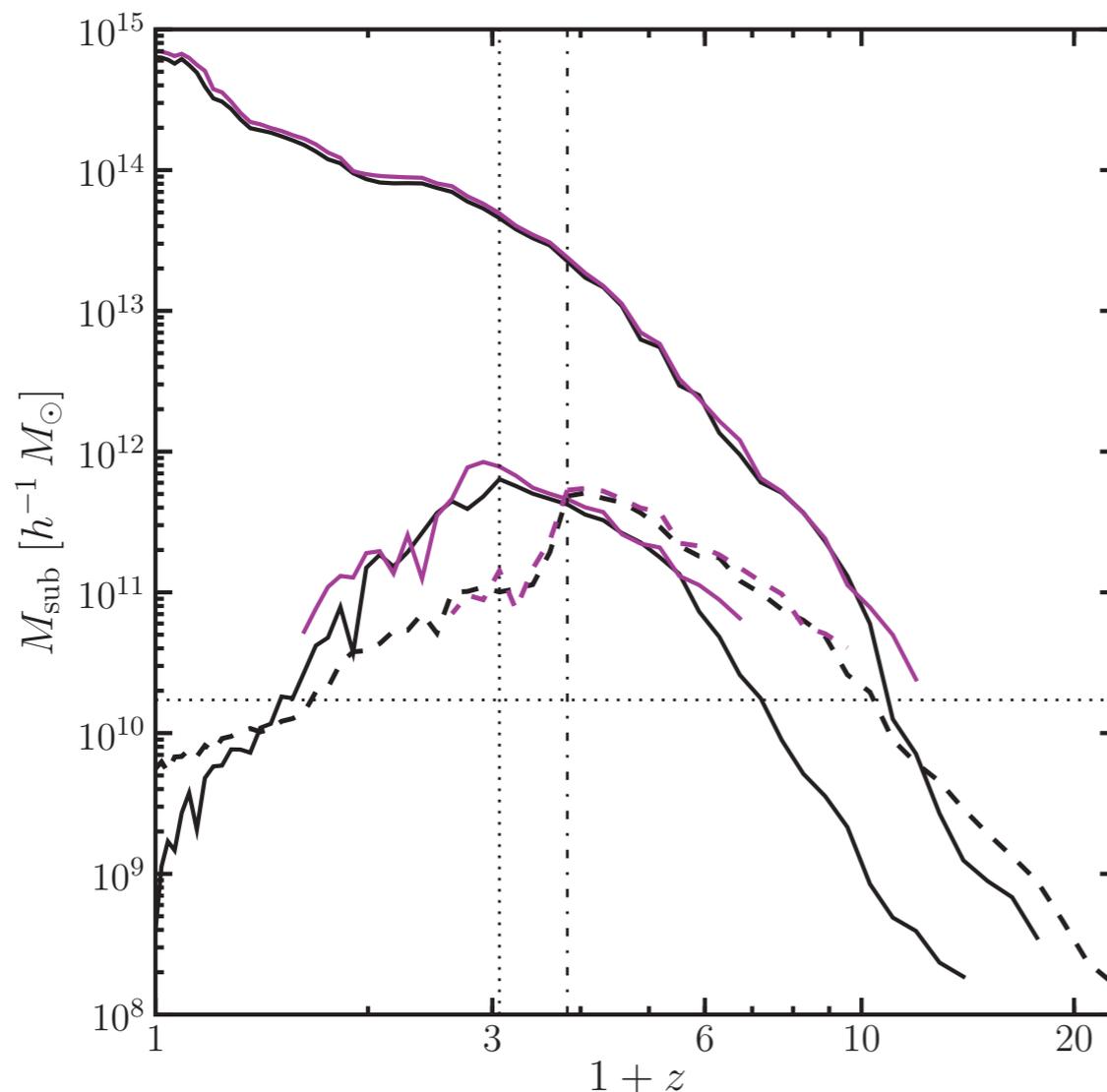


halos selected from the Millennium-II Simulation

SDSS DR7: satellites of isolated galaxies

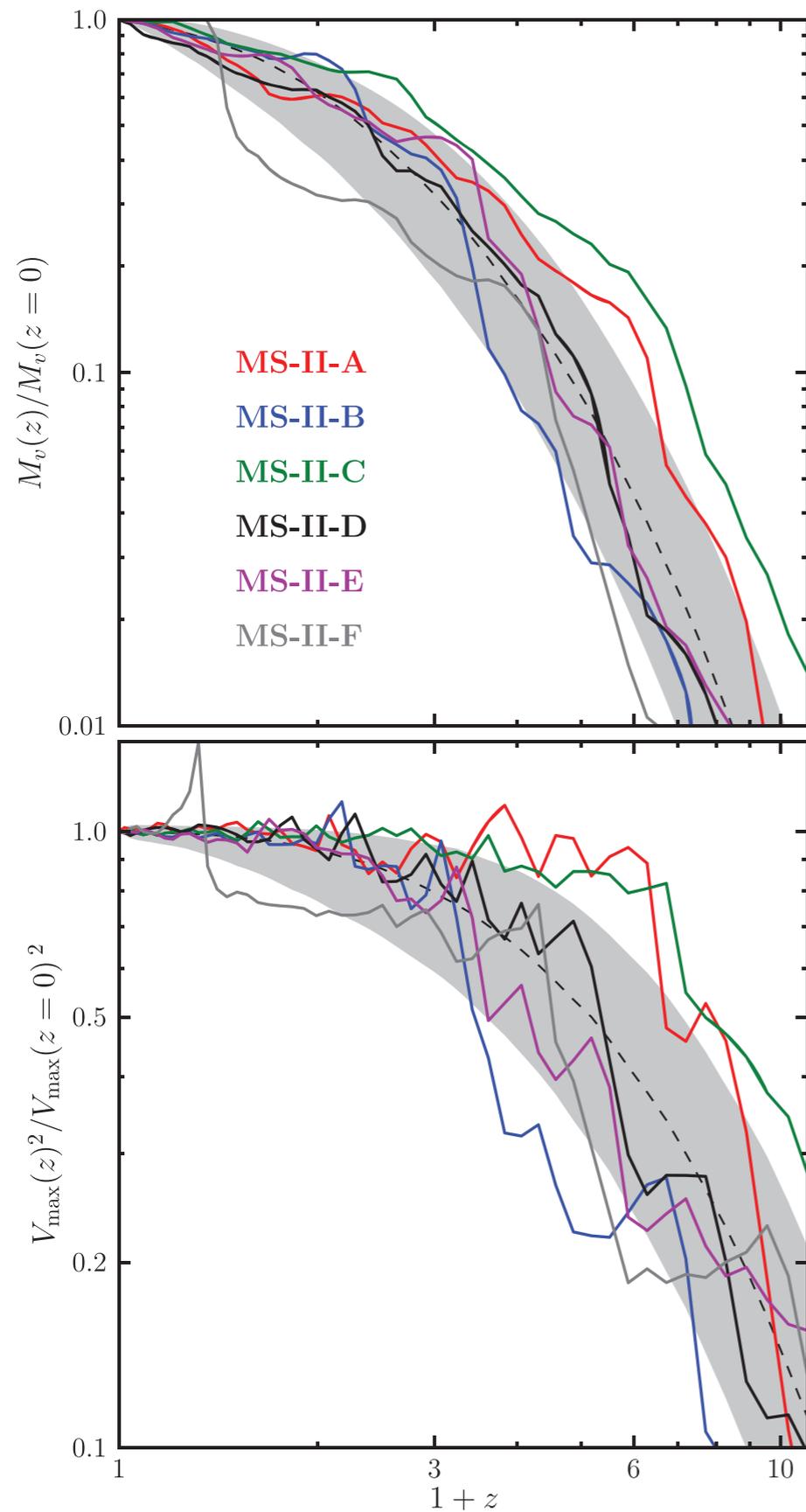
How I learned to love SQL

- Millennium II: very highly clustered
 - ▶ largest tree file in MS-II has 90 million halos (compare to 500,000 in Millennium) - 1/6 of all halos in the simulation (!!)
 - ▶ accessing the formation history of an individual halo is often very inefficient



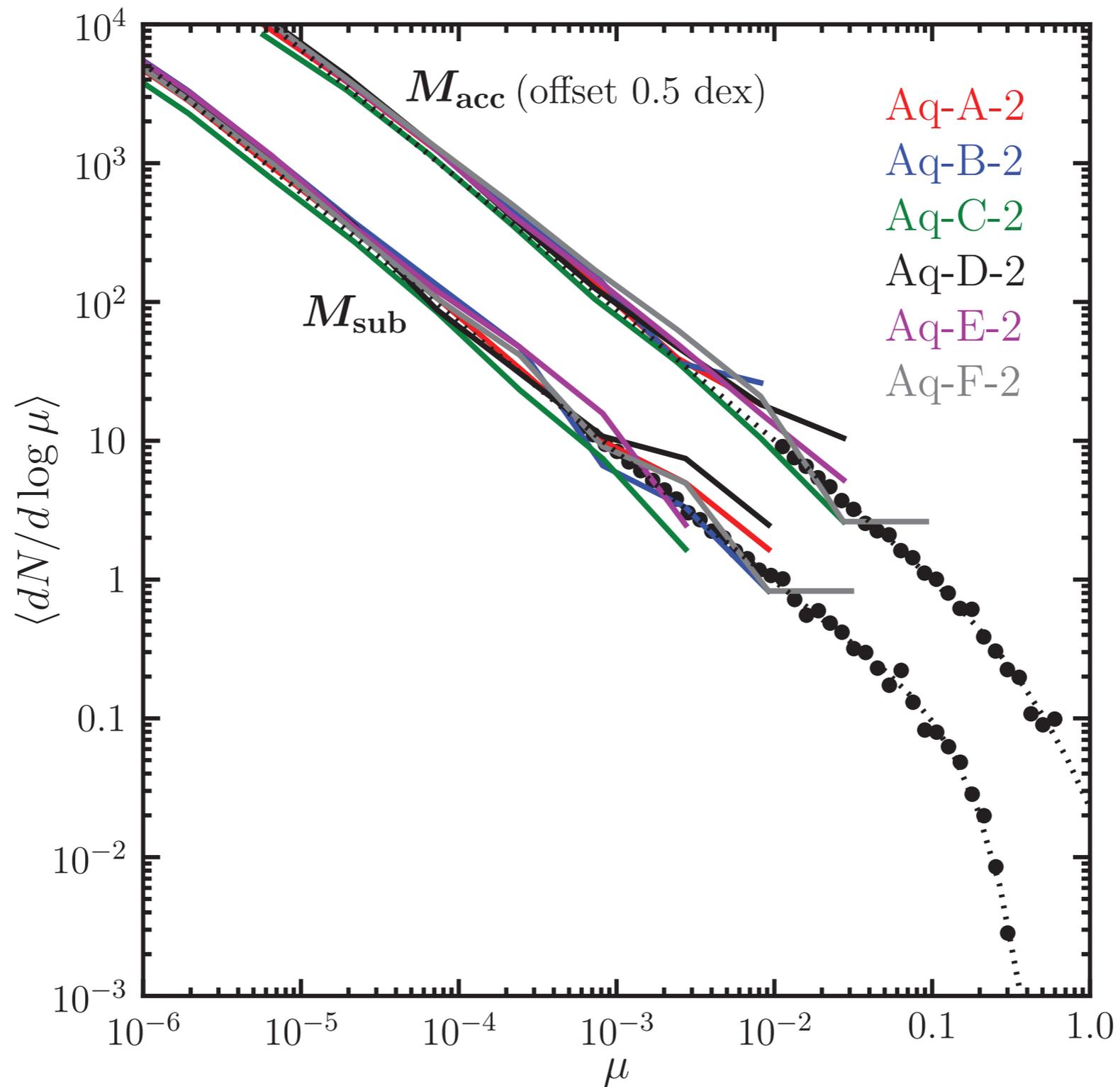
Making this plot requires reading and holding 10 GB of data in memory if using data files; requires ~1000 bytes when using the database

Mass accretion histories for thousands of Milky Way-mass halos



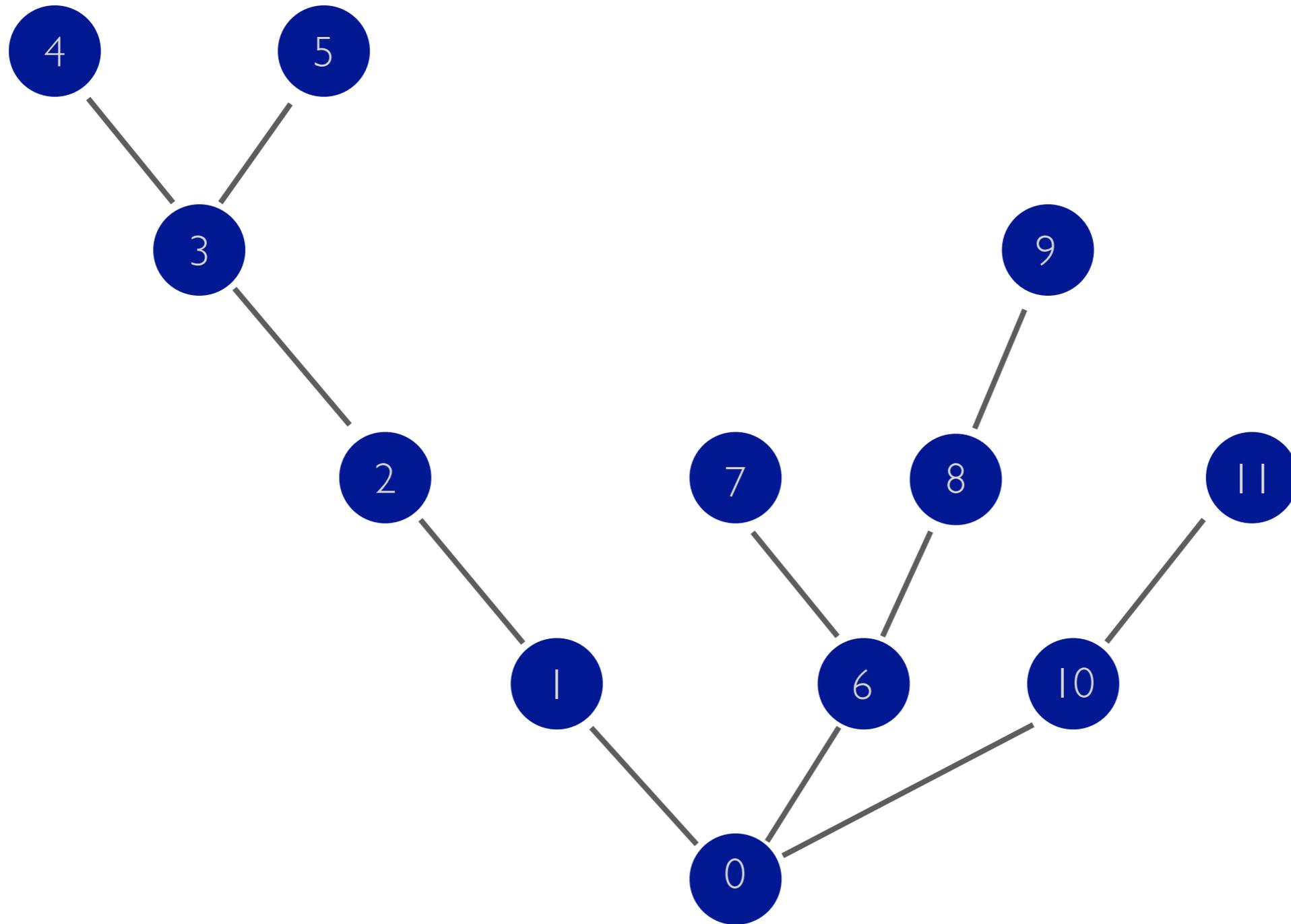
Dashed: median mass accretion history
Shaded: 68% of data

Binned subhalo mass functions

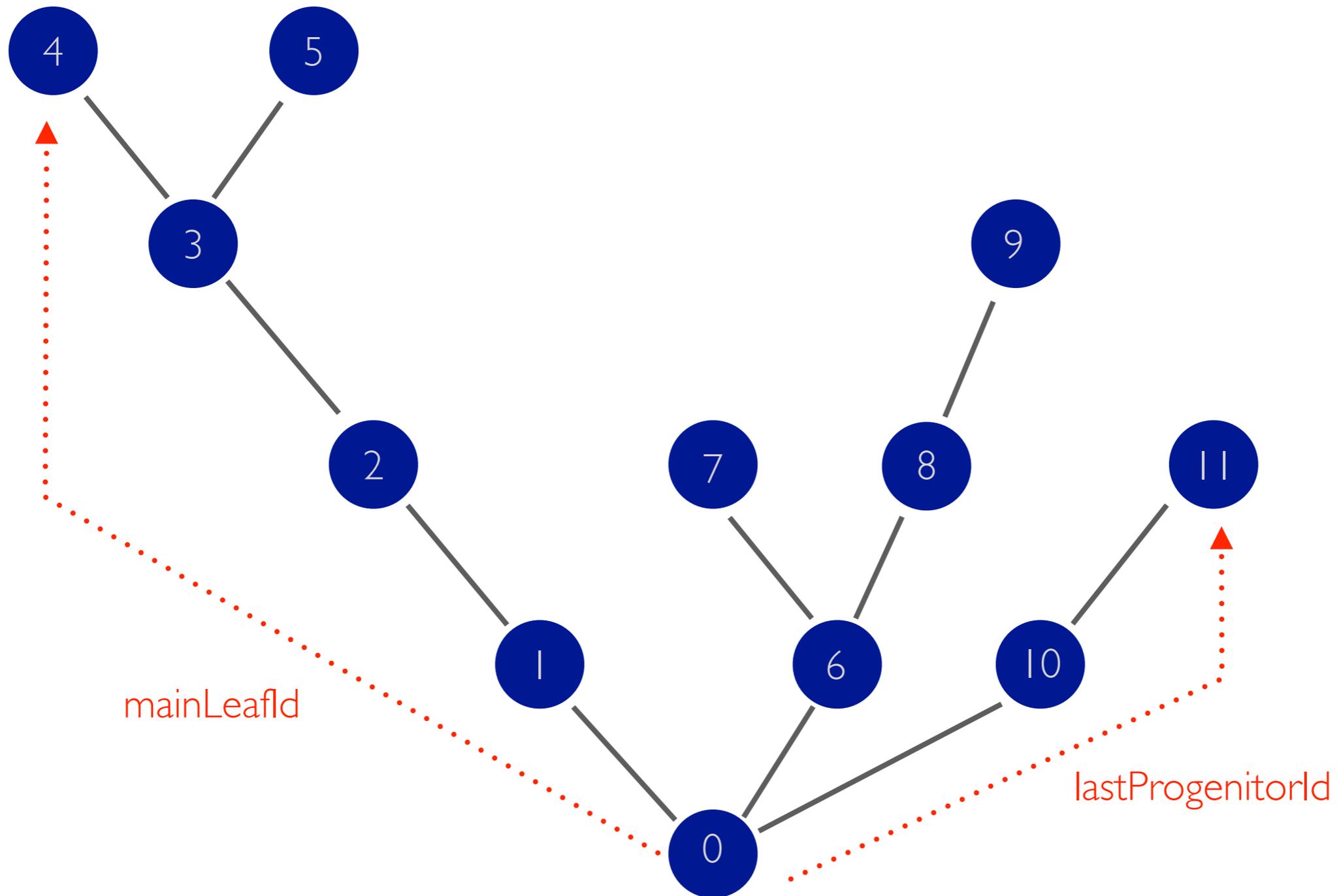


MBK, Springel, White, Jenkins (2010)

Additions to tree pointers from MS to MS-II



Additions to tree pointers from MS to MS-II



Functions and functionality I learned to love

- insert into
 - personal databases allow very customizable data storage and access
- rank () over (partition by)
 - use this often to find halo properties at a specific time (e.g., when V_{\max} or mass is maximized) along a main progenitor branch for ensembles of halos
- row_number() over
 - rank objects at different redshifts, then compare easily across redshifts
- create your own function
 - example: distance between two points, taking into account periodic boundaries

```
CREATE FUNCTION distance
(
  @boxsize FLOAT,
  @x1 FLOAT, @y1 FLOAT, @z1 FLOAT,
  @x2 FLOAT, @y2 FLOAT, @z2 FLOAT
)
RETURNS FLOAT AS
BEGIN
  RETURN
  POWER(
    POWER(ABS(@x1-@x2)-@boxsize*FLOOR(abs(@x1-@x2)/@boxsize*2.0),2.0)+
    POWER(ABS(@y1-@y2)-@boxsize*FLOOR(abs(@y1-@y2)/@boxsize*2.0),2.0)+
    POWER(ABS(@z1-@z2)-@boxsize*FLOOR(abs(@z1-@z2)/@boxsize*2.0),2.0)
  ,0.5)
```

Millennium-II Papers

[Home](#)[Data](#)[Documentation](#)[Papers](#)[Links](#)[Private](#)

Reference papers:

- Lemson G. and the Virgo Consortium (2006), [arXiv:astro-ph/0608019](#) (public release of the Millennium Simulation database)
- Springel V., et al. (2005), [Nature](#), **435**, 629 (main paper for the Millennium Simulation)
- Springel V., et al. (2008), [MNRAS](#), **391**, 1685 (introductory paper for the Aquarius Project)
- Springel V., White S. D. M., Tormen G., Kauffmann G. (2001), [MNRAS](#), **328**, 726 (introduction of the SUBFIND algorithm)

Papers using data from the Millennium-II Simulation:

Please contact us if you know of any papers that are missing from this list of papers that directly use data from the Millennium-II Simulation.

1. [arXiv:0903.3041](#) [[pdf](#), [other](#)]
Title: Resolving Cosmic Structure Formation with the Millennium-II Simulation
Authors: [Michael Boylan-Kolchin](#) (1), [Volker Springel](#) (1), [Simon D. M. White](#) (1), [Adrian Jenkins](#) (2), [Gerard Lemson](#) (3, 4)
(1) MPA, (2) Durham, (3) ARI-ZAH, (4) MPE
Comments: 16 pages, 13 figures; matches version published in MNRAS. Halo catalogs and merger trees, along with movies, images, and additional information, are available at [this http URL](#).
Journal-ref: Mon.Not.Roy.Astron.Soc. 398, 1150 (2009)
2. [arXiv:0908.2428](#) [[ps](#), [pdf](#), [other](#)]
Title: Extragalactic gamma-ray background radiation from dark matter annihilation
Authors: [Jesus Zavala](#), [Volker Springel](#), [Michael Boylan-Kolchin](#)
Comments: 21 pages, 16 figures, submitted to MNRAS

Sixty papers have directly used MS-II data

58. [arXiv:1210.7185](#) [[pdf](#), [ps](#), [other](#)]
Title: Modeling the Redshift Evolution of the Normal Galaxy X-ray Luminosity Function
Authors: [M. Tremmel](#), [T. Fragos](#), [B. D. Lehmer](#), [P. Tzanavaris](#), [K. Belczynski](#), [V. Kalogera](#), [A. R. Basu-Zych](#), [W. M. Farr](#), [A. Hornschemeier](#), [L. Jenkins](#), [A. Ptak](#), [A. Zezas](#)
Comments: Submitted to ApJ, 16 pages, 3 tables, 7 figures
59. [arXiv:1211.3752](#) [[pdf](#), [ps](#), [other](#)]
Title: Massive black hole seeds born via direct gas collapse in galaxy mergers: their properties, statistics and environment
Authors: [Silvia Bonoli](#), [Lucio Mayer](#), [Simone Callegari](#) (University of Zurich)
Comments: 14 pages, 10 figures, submitted to MNRAS
60. [arXiv:1212.1717](#) [[pdf](#), [ps](#), [other](#)]
Title: Simulations of the galaxy population constrained by observations from $z=3$ to the present day: implications for galactic winds and the fate of their ejecta
Authors: [Bruno Henriques](#) (MPA), [Simon White](#) (MPA), [Peter Thomas](#) (Sussex), [Raul Angulo](#) (Stanford), [Qi Guo](#) (Beijing), [Gerard Lemson](#) (MPA), [Volker Springel](#) (HITS)
Comments: 25 pages, 14 figures, submitted to MNRAS

Thank you to everyone involved!