GCE in...



Rob Yates - L-Galaxies discussion - 28th November 2018

GCE set-up

• <u>Same implementation as in Yates+13:</u>

- Delayed enrichment from SNe-II, SNe-Ia, and AGB stars
- Mass- and metallicity-dependent yields & stellar lifetimes (Portinari+98; Thielemann+03; Marigo+01)
- Constant Chabrier+03 IMF
- Power-law SN-la DTD (Maoz+12)
- Store SF and metallicity histories
- SN feedback occurs when stars die

- Main changes:
 - Percentage of stars that form as SN-Ia progenitors: Before: A' = 0.108% Now: A' = 0.154%
 - Fraction of SN ejecta directly enriching hot gas: Before: $f_{SNII} = f_{SNII} = 0.0$ Now: $f_{SNII} = f_{SNII} = 0.3$
 - SN energy parameter: Before: 2.5x10⁻⁵¹ erg

Now: *5x10⁻⁵¹ erg*



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General galaxy properties



 Galaxy colours are systematically bluer than in the Henriques+15 version of L-Galaxies



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• Star formation is a little lower at high redshifts





- MZ_{*}R has similar slope to recent measurements for SDSS galaxies (Zahid+17)
 - Normalisation is a little low (perhaps expected, given bluer colours)
- However, model Z_{*} is mass weighted, whereas observed Z_{*} is V-band-luminosity weighted

Cold Gas: MZ_qR





 <u>SFR-weighted</u> MZ_gR is closer to (my favourite) observed MZ_gR from strong-lines (orange)



 <u>Mass-weighted</u> MZ_gR is closer to (my favourite) observed MZ_gR from electron temperatures (green)

NB: Normalisation of the observed MZ_gR varies by up to ~0.6 dex. Difficult to use as a constraint



- Iron abundance in model groups & clusters is in reasonable agreement with observations
- Improvements in X-ray spectral fitting and radial Fe gradient measurements for groups have brought observations in to line with models (Mernier+17,18)

Follow-up paper on GCE

- Main model paper (Henriques+):
 - Present key redshift O scaling relations (MZ_{*}R, MZ_gR, TZ_{Fe}R)
 - Present Z_{*} and Z_g radial gradients in star-forming galaxies in stellar mass bins



- Follow-up GCE paper (Yates+):
 - Evolution of the MZ_gR relation
 - Evolution of metallicity gradients in galaxies. Dependencies on SFH and merger history, etc
 - Chemical composition in low-redshift ETGs
 - Chemical composition in the MW stellar disc (compare to APOGEE, AMBRE, GAIA-ESO, etc)
 - Consider changes to the GCE set-up: ISM-density dependent direct ejection, constrain *A*' and *f*_{SN} using MCMC, etc...



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