Galaxy formation with L-GALAXIES: Catalogue Description

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In this supplementary material, we describe the full catalogues of galaxy merger trees that are produced by the Ayromlou et al. 2021 version of the L-GALAXIES semi-analytical model. A full description of the galaxy formation model used to produce these catalogues can be found here. In addition, the main paper describing this model is given here. Please address your questions about this model/catalogue to Reza Ayromlou (ayromlou@mpa-garching.mpg.de)

| Quantity [unit] | Shape | Description |
|--|-------|---|
| Туре | N | Indicates whether the galaxy is a "central" (at the center of its FOF group, type=0), a "satellite" (within its own subhalo but not at the center of its FOF group, type=1), or an "orphan" (a satellite that has lost its subhalo, type=2) |
| HaloIndex | N | The unique identifier of the subhalo |
| SnapNum | N | The snapshot number where this galaxy was identified |
| LookBackTimeToSnap [year] | N | The look back time (in years) from $z = 0$ to the redshift of the galaxy |
| $\begin{array}{c} Central_M_Crit200\\ [10^{10}M_{\odot}/h] \end{array}$ | N | The virial mass $(M_{200,crit})$ of the FOF group the galaxy resides in |
| Central_R_Crit200 [Mpc/h] | N | The virial radius $(R_{200,crit})$ of the FOF group the galaxy resides in |
| Pos [cMpc/h] | N,3 | The galaxy's position in comoving coordinates |
| Vel [km/s] | N,3 | The galaxy's velocity |
| SuhaloLen | N | Number of particles associated with the subhalo hosting this galaxy |
| $\rm M_Crit200~[10^{10} M_{\odot}/h]$ | N | Virial mass $(M_{200,crit})$ of the subhalo this galaxy was in when it was last a type 0 galaxy. I.e. current virial mass for type 0 galaxies, "infall virial mass" for type 1,2 galaxies |
| $R_{-}Crit200 \ [Mpc/h]$ | N | Virial radius $(R_{200,crit})$ of the subhalo this galaxy was in when it was last a type 0 galaxy. I.e. current virial radius for type 0 galaxies, "infall virial radius" for type 1,2 galaxies |
| Vvir [km/s] | N | Virial velocity of the subhalo this galaxy was in when it was last a type 0 galaxy. I.e. current virial velocity for type 0 galaxies, "infall virial velocity" for type 1,2 galaxies |
| Vmax [km/s] | N | Maximum rotational velocity of the subhalo of this galaxy. This property continues to be updated even after the galaxy becomes a type 1 |
| $\begin{array}{l} {\rm HaloSpin} \\ [({\rm Mpc/h})({\rm km/s})] \end{array}$ | N,3 | the spin of the cold gas disk |

Table 1: Main Properties of Galaxies

| InfallVmax [km/s] | N | Maximum rotational velocity of the subhalo of this galaxy. This property continues to be updated even after the galaxy becomes a type 1 |
|--|---|--|
| InfallVmaxPeak [km/s] | N | Maximum past rotational velocity of the subhalo of this galaxy |
| InfallSnap | Ν | Most recent (largest) snapnum at which this galaxy's type changed from 0 to 1 or 2 $$ |
| InfallHotGasMass $[10^{10}M_{\odot}/h]$ | Ν | Mass in hot gas at the time of infall (same as hot Gas for type 0 galaxies) $% \left({{\left[{{{\left[{{\left[{{\left[{{\left[{{\left[{{\left[$ |
| HotGasRadius [Mpc/h] | Ν | Radius out to which hot gas extends: Rvir for type 0; 0 for type 2; maximum radius out to which hot gas is not stripped for type 1. |
| OriMergTime [year] | N | Estimated dynamical friction time (in years) when the merger clock was set. Only calculated for type 2 galaxies. |
| MergTime [year] | Ν | Estimated remaining merging time (in years). OriMergeTime - time since the merger clock is set. Only calculated for type 2 galaxies |
| m ColdGasMass $ m [10^{10}M_{\odot}/h]$ | N | Mass in the cold gas disk |
| H2fraction | Ν | Fraction of cold gas mass that is in H2 |
| Stellar Mass $[10^{10} {\rm M}_\odot/{\rm h}]$ | Ν | Total mass in stars in the disk and the bulge together |
| StellarDiskMass $[10^{10}M_{\odot}/h]$ | Ν | Mass of stars in the disk |
| $\begin{array}{l} StellarBulgeMass \\ [10^{10}M_{\odot}/h] \end{array}$ | Ν | Mass of stars in the bulge |
| HotGasMass $[10^{10}M_{\odot}/h]$ | Ν | Mass in hot gas |
| Ejected Mass $[10^{10} \rm M_\odot/h]$ | Ν | Mass in the ejected gas component |
| BlackHoleMass $[10^{10}M_{\odot}/h]$ | Ν | Mass of the central black hole |
| $\begin{array}{l} HaloStellarMass \\ [10^{10}M_{\odot}/h] \end{array}$ | Ν | Mass in intra-cluster stars |
| $\begin{array}{l} {\rm PrimordialAccretionRate} \\ {\rm [M_{\odot}/yr]} \end{array}$ | N | Accretion rate of primordial gas |
| CoolingRadius [Mpc/h] | Ν | The radius within which the cooling time scale is shorter than the dy- namical timescale |
| CoolingRate $[M_{\odot}/yr]$ | Ν | Cooling rate |
| $\begin{array}{l} {\rm CoolingRate_beforeAGN} \\ {\rm [M_{\odot}/yr]} \end{array}$ | Ν | Cooling rate if there was no AGN feedback |
| $\begin{array}{l} QuasarAccretionRate \\ [M_{\odot}/yr] \end{array}$ | Ν | Rate at which cold gas is accreted into the central black hole in the quasar mode. |
| $\begin{array}{l} {\rm RadioAccretionRate} \\ {\rm [M_{\odot}/yr]} \end{array}$ | Ν | Rate at which hot gas is accreted into the central black hole in the radio mode |
| StarFormationRate $[M_{\odot}/yr]$ | N | Star formation rate |
| StarFormationRateBulge $[M_{\odot}/yr]$ | N | Star formation rate in bulge |
| XrayLum $[\log_{10}(\text{erg/sec})]$ | N | Log10 of X-Ray luminosity in erg/sec |
| BulgeSize [Mpc/h] | N | Half mass radius of bulge |

| StellarDiskRadius [Mpc/h] | N | Size of the stellar disk, 3x the scale length |
|---|------|---|
| GasDiskRadius [Mpc/h] | Ν | Size of the cold gas disk |
| StellarHalfMassRadius [Mpc/h] | N | Half-mass radius of the stellar disk |
| $\begin{array}{l} StellarHalfLightRadius \\ [Mpc/h] \end{array}$ | Ν | Stellar half light radius |
| CosInclination [deg] | Ν | Inclination of the galaxy. Derived from the angle between the total and z-axis stellar spins of the galaxy |
| DisruptOn | Ν | 0: galaxy merged onto merger center; 1: galaxy was disrupted before merging onto its descendant, matter went into ICM of merger center |
| MergeOn | N | 0: merger clock not set yet; 1: type 1 galaxy with baryon mass ; halo mass, separate dynamical friction time calculated; 2: this galaxy is type 2 and will merge into the merger center in the next snapshot; 3: this galaxy is type 1 and will merge into the central galaxy of the main halo in the next snapshot |
| MagDust | N,20 | Rest-frame absolute magnitude of the galaxy (dust extinction included). Description of the columns is given below |
| Mag | N,20 | Rest-frame absolute magnitude of the galaxy. Description of the columns is given below |
| MagBulge | N,20 | Rest-frame absolute magnitude of the galaxy's bulge. Description of the columns is given below |
| MassWeightAge $[10^9 \mathrm{yr}]$ | Ν | The age of this galaxy, weighted by mass of its components |
| rBandWeightAge $[10^9 \mathrm{yr}]$ | Ν | The age of this galaxy, weighted by it's uncorrected r-band magnitude |
| $\begin{array}{c} \mathrm{rho_LBE_Gas} \\ \mathrm{[10^{10}M_{\odot}h^2/Mpc^3]} \end{array}$ | N | Local background environment density of gas |
| $\frac{\rm rho_LBE_Total}{[10^{10}M_{\odot}h^2/Mpc^3]}$ | Ν | Local background environment density |
| Vel_LBE [km/s] | N | Local background environment velocity |

Table 2: Additional Properties of Galaxies

| Quantity [unit] | Shape | Description |
|--|-------|--|
| m ColdGasMassRings $[10^{10}M_{\odot}/h]$ | N,12 | The unique identifier of each ring in a galaxy. Runs from 0 to 11, where ring 0 is the innermost ring |
| H2fractionRings | N,12 | Fraction of cold gas mass in each ring that is in H2 |
| $\begin{array}{l} StellarDiskMassRings \\ [10^{10}M_{\odot}/h] \end{array}$ | N,12 | Mass of stars in each ring of the disk |
| $\begin{array}{l} StellarBulgeMassRings \\ [10^{10}M_{\odot}/h] \end{array}$ | N,12 | Mass of stars in each ring of the bulge |
| StellarMassFromInSitu $[10^{10}M_{\odot}/h]$ | N | Stellar mass formed in-situ in the galaxy, rather than accreted or formed in merger-induced starbursts. I.e. total mass of stars formed secularly in progenitors along this galaxy's main progenitor branch. |

| StellarMassFromMergers $[10^{10}M_{\odot}/h]$ | N | Stellar mass accreted onto the galaxy in mergers, rather than formed in-situ or via merger-induced starbursts. I.e. total mass of stars formed in progenitors not along this galaxy's main progen- itor branch |
|--|---------|---|
| StellarMassFromBursts $[10^{10}M_{\odot}/h]$ | N | Mass formed in merger-induced starbursts in the galaxy, rather than accreted in mergers of formed secularly. |
| $\frac{MetalsColdGasMass}{[10^{10}M_{\odot}/h]}$ | N,3 | Mass in metals in the cold gas disk |
| $\begin{array}{c} MetalsColdGasMassRings \\ [10^{10}M_{\odot}/h] \end{array}$ | N,12,3 | Mass of metals in each ring of the cold gas disk |
| $[10^{10} M_{\odot}/h]$ | N,3 | Mass in metals in stars in the disk and the bulge together |
| $\begin{array}{l} MetalsStellarDiskMass \\ [10^{10}M_{\odot}/h] \end{array}$ | N,3 | Mass in metals in stars in the disk |
| $\begin{array}{l} MetalsStellarBulgeMass \\ [10^{10}M_{\odot}/h] \end{array}$ | N,3 | Mass in metals in stars in the bulge |
| MetalsStellarDiskMassRings $[10^{10}M_{\odot}/h]$ | N,12,3 | Mass of metals in each ring of the stellar disk |
| $\label{eq:MetalsStellarBulgeMassRings} \end{tabularBulgeMassRings} \end{tabularBulgeMassRings} \end{tabularBulgeMassRings} \end{tabularBulgeMassRings} \end{tabularBulgeMassRings}$ | N,12,3 | Mass of metals in each ring of the stellar bulge |
| $\begin{array}{l} MetalsHotGasMass\\ [10^{10}M_{\odot}/h] \end{array}$ | N,3 | Mass in metals in the hot gas |
| $\begin{array}{l} MetalsEjectedMass \\ [10^{10}M_{\odot}/h] \end{array}$ | N,3 | Mass of metals in the ejected mass component |
| $\begin{array}{l} MetalsHaloStellarMass\\ [10^{10}M_{\odot}/h] \end{array}$ | N,3 | Mass of metals in halo stars |
| $\begin{array}{l} StarFormationRateRings\\ [M_{\odot}/yr] \end{array}$ | N,12 | Average star formation rate across the snapshot in each ring |
| DiskMass_elements $[M_{\odot}]$ | N,11 | Mass of elements in the stellar disk |
| BulgeMass_elements $[M_{\odot}]$ | N,11 | Mass of elements in the stellar bulge |
| $DiskMassRings_elements \\ [M_{\odot}]$ | N,12,11 | Mass of elements in each ring of the stellar disk |
| $\frac{\rm BulgeMassRings_elements}{[M_{\odot}]}$ | N,12,11 | Mass of elements in each ring of the stellar bulge |
| ColdGas_elements $[M_{\odot}]$ | N,11 | Mass of elements in the cold gas disk |
| ColdGasRings_elements $[{\rm M}_{\odot}]$ | N,12,11 | Mass of elements in each ring of the cold gas disk |
| HotGas_elements $[M_{\odot}]$ | N,11 | Mass of elements in the hot gas |
| ICM_elements $[M_{\odot}]$ | N,11 | Mass of elements in halo stars |
| Ejected Mass_elements $[{\rm M}_{\odot}]$ | N,11 | Mass of elements in the ejected mass component |

^{*} For each "element" dataset, eleven elements are included: H, He, C, N, O, Ne, Mg, Si, S, Ca, Fe * For each "ring" dataset, the outer ring radius in kpc is given by the following list:

^[0.029, 0.059, 0.12, 0.24, 0.47, 0.94, 1.9, 3.8, 7.5, 15, 30, 60]

Table 3: Magnitudes

For each of the three Mag^{*} datasets, all magnitudes are rest-frame, absolute (AB). The twenty entries correspond to (in order):

| Entry | Band |
|-------|--|
| 0 | Rest-frame absolute (AB) magnitude in the Johnson-Bessel U filter ($\lambda = 0.36\mu$ m) of the galaxy |
| 1 | Rest-frame absolute (AB) magnitude in the Johnson-Bessel B filter ($\lambda = 0.435 \mu$ m) of the galaxy |
| 2 | Rest-frame absolute (AB) magnitude in the Johnson-Bessel V filter ($\lambda = 0.55 \mu$ m) of the galaxy |
| 3 | Rest-frame absolute (AB) magnitude in the Cousins Rc filter ($\lambda = 0.64 \mu m$) of the galaxy |
| 4 | Rest-frame absolute (AB) magnitude in the Cousins Ic filter ($\lambda = 0.79 \mu m$) of the galaxy |
| 5 | Rest-frame absolute (AB) magnitude in the VISTA Z filter ($\lambda = 0.88 \mu$ m) of the galaxy |
| 6 | Rest-frame absolute (AB) magnitude in the VISTA Y filter ($\lambda = 1.02 \mu m$) of the galaxy |
| 7 | Rest-frame absolute (AB) magnitude in the VISTA/2MASS J filter ($\lambda = 1.26 \mu m$) of the galaxy |
| 8 | Rest-frame absolute (AB) magnitude in the VISTA/2MASS H filter ($\lambda = 1.60 \mu m$) of the galaxy |
| 9 | Rest-frame absolute (AB) magnitude in the Johnson-Bessel K ($\lambda = 2.22 \mu m$) filter of the galaxy |
| 10 | Rest-frame absolute (AB) magnitude in the VISTA/2MASS Ks ($\lambda = 2.16 \mu m$) filter of the galaxy |
| 11 | Rest-frame absolute (AB) magnitude in the IRAC 3.6um filter of the galaxy |
| 12 | Rest-frame absolute (AB) magnitude in the IRAC 4.5um filter of the galaxy |
| 13 | Rest-frame absolute (AB) magnitude in the IRAC 5.8um filter of the galaxy |
| 14 | Rest-frame absolute (AB) magnitude in the IRAC 8.0um filter of the galaxy |
| 15 | Rest-frame absolute (AB) magnitude in the SDSS u filter ($\lambda = 0.355 \mu$ m) of the galaxy |
| 16 | Rest-frame absolute (AB) magnitude in the SDSS g filter ($\lambda = 0.469 \mu$ m) of the galaxy |
| 17 | Rest-frame absolute (AB) magnitude in the SDSS r filter ($\lambda = 0.617 \mu m$) of the galaxy |
| 18 | Rest-frame absolute (AB) magnitude in the SDSS i filter ($\lambda = 0.748 \mu m$) of the galaxy |
| 19 | Rest-frame absolute (AB) magnitude in the SDSS z filter ($\lambda = 0.893 \mu$ m) of the galaxy |